

# Fractons in effective field theories for spontaneously broken translations

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Based on:

- R. Argurio, C. Hoyos, D. Musso, D. N. *Phys. Rev. D* **104** no. 10, (2021) 105001, [arXiv:2107.03073 \[hep-th\]](https://arxiv.org/abs/2107.03073).

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# Structure of the talk

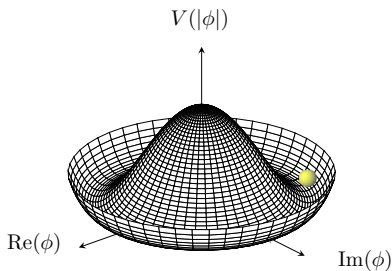
- 1 Goldstone Physics
- 2 Fractons in field theory
- 3 Fractons in EFTs for spontaneously broken translations
- 4 Conclusion: results and outlooks

# Key concepts

- We consider field theories defined on Minkowski spacetime at  $T = 0$  &  $\mu = 0$ , with  $d \geq 2$ . We chose

$$\text{diag}(\eta) = (+, -, \dots, -) \quad ; \quad \hbar = 1 = c .$$

- Symmetry: the EOM are unchanged by the considered transformation  $\Leftrightarrow \delta S = 0$ .
- Spontaneous symmetry breaking (SSB): a stable state of the system transforms under certain symmetries of the theory.



# Goldstone's theorem

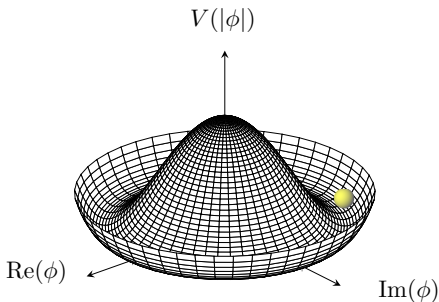
## Goldstone's theorem [Goldstone, Nuovo Cim. '61]

If a theory has a global continuous symmetry group  $G$  spontaneously broken to a subgroup  $H$ . Then, the spectrum of the theory will contain at least one gapless mode ( $\omega \xrightarrow{\vec{k} \rightarrow \vec{0}} 0$ ): Nambu Goldstone mode (NG mode).

- Let  $v$  be the VEV and  $X$  be a broken generator.

NG candidate:  $e^{i\pi(x)X} v$

- NG candidates are weakly coupled in the IR.



# More on the NG modes

## Perturbation theory around the VEV:

- NG candidates weakly coupled in the IR  $\Rightarrow \boxed{\mathcal{L}_{\text{int}} \sim \dots \partial\pi \dots}$ .
- Spontaneous breaking: **symmetry** still there but **non-linearly realized**.  
Ex. for  $U(1)$ -SSB:

$$\begin{aligned} (v + \rho(x))e^{i\pi(x)} &\xrightarrow{U(1)} (v + \rho(x))e^{i(\pi(x)+\alpha)} \\ &\Rightarrow \pi(x) \xrightarrow{U(1)} \pi(x) + \alpha . \end{aligned}$$

NG candidates transform "shift like"

- Parametrization of the NG candidates ( $\sim e^{i\pi(x)X} v$ )  
 $\neq$  parametrization diagonalizing the kinetic matrix  
 $\rightarrow$  involved identification of the modes.  
 $\rightarrow$  Ward-Takahashi Identities can help.

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## Motivations:

- Symmetries & SSB  $\Rightarrow$  universal exact results.
- **Hastily said**, Goldstone physics universally describes IR physics.

$\Rightarrow$  Light mesons physics, phase transitions, stellar superfluids . . .



# Motivations and Research Directions

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## Research Directions:

- **Counting rules.**
- Conditions to have SSB (ex: Coleman's theorem).
- How do NG modes interact with non-symmetry originated particles ?

# Counting rule for internal symmetries

- $\exists$  complete counting rule for the breaking of compact internal symmetries. [Watanabe et al., PRL. '12]
- Has been obtained through the most general EFT for  $G \rightarrow H$ :

$$\mathcal{L}_{\text{eff}}(\pi) = c_a(\pi)\partial_t\pi^a + \frac{1}{2}g_{ab}(\pi)\partial_t\pi^a\partial_t\pi^b - \frac{1}{2}\bar{g}_{ab}(\pi)\partial_i\pi^a\partial_i\pi^b + \mathcal{O}(\partial^n).$$

- The reduction of the number of NG modes is purely dynamical & is due to  $c_a(\pi)\partial_t\pi^a$ :

$$\mathcal{L} \sim \pi^1\partial_t\pi^2 - \pi^2\partial_t\pi^1 + \dots.$$

$$P_a \equiv \frac{\partial\mathcal{L}}{\partial(\partial_t\pi^a)} \rightsquigarrow \text{canonically conjugated pairs } (\pi^1, -\pi^2) \& (\pi^2, \pi^1).$$

- For relativistic theories:  $c_a(\pi) = 0 \Rightarrow n_{\text{NG}} = n_{\text{BS}}$ .

# Counting rule for spacetime symmetries

## Why is it more complicated ?

- Non-compact symmetries  $\rightsquigarrow$  less “nice” mathematical properties.
- The VEV is spacetime dependent  $\rightsquigarrow$  functional aspect of QFT is emphasized (minimization of energy etc.).
- Effective Lagrangians less constrained  $\rightsquigarrow$  more complicated dispersion relation.
- Derrick's theorem suggests the **need of higher derivative terms** in the fundamental theory  $\rightsquigarrow$  even toy models are complicated!

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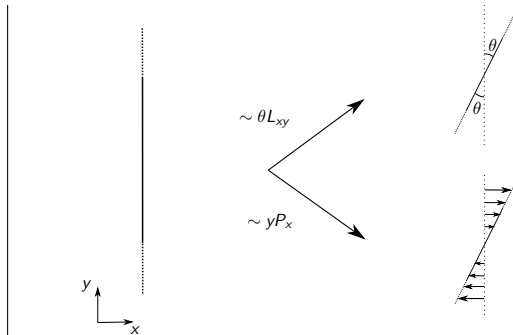
## But does it affect the counting rules ?

- Yes! Ex: breaking of the conformal group to Poincaré group  $\rightsquigarrow$  1 NG mode for  $d + 1$  broken generators.

# Counting rule for spacetime symmetries

Some clues for the reduction of NG modes:

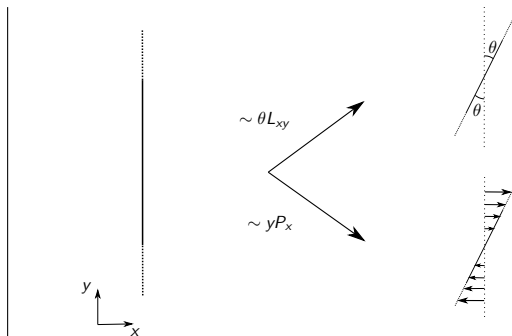
Non-dynamical reduction due to definitions and trsf. laws.



# Counting rule for spacetime symmetries

Some clues for the reduction of NG modes:

Non-dynamical reduction due to definitions and trsf. laws.



Formalized through the **Inverse Higgs Constraints (IHC)**

If  $[P_\mu, X_1] \sim X_2$   $\Rightarrow$  possible to write  $\mathcal{L}_{\text{eff}}(\pi_1)$  w/o  $\pi_2$ ,

where  $P_\mu$  an unbroken translation gen. and  $X_i$  are broken gen.

Under which criteria impose IHC ??

[Ivanov et al., Teo. Mat. Fiz. '75]

## What to bear in mind:

- SSB  $\Rightarrow$  NG candidates  $\pi_a \Rightarrow$  independent massless NG modes.
- $\pi_a$  transform “shift like”.
- $\mathcal{L}_{\text{int}}(\pi) \sim \dots \partial\pi \dots$
- Counting rule for internal SSB but not for spacetime SSB.
- **Possibly** NG reduction through  $[P_\mu, X_1] \sim X_2$  (IHC).
- Translation SSB  $\rightsquigarrow$  higher derivative terms.

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# Introduction to fractons

## Vocabulary:

- Fractonic modes = excitations with reduced mobility.
- Fractons = immobile modes.
- Subdimensional modes = modes which can propagate in a restricted number of directions.

## Historically:

- First noticed in lattice models. [Vijay et al., PRB. '15]
- Then extended to continuous model via (spatial) higher rank gauge theories ( $A_{ij}$ ). [Pretko, PRB. '17]

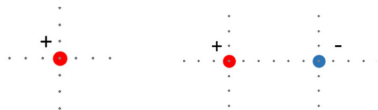
# Introduction to fractons

## Intuition:

- Gauge theories  $\Rightarrow$  gauge constraints in the Hamiltonian formalism (ex: Gauss law).
- Gauge constr. interpreted as conservation laws for external sources.
- Higher rank gauge theo.  $\Rightarrow$  more cons. laws  $\Rightarrow$  restricted dynamics.

$$\partial_i \partial_j E^{ij} = \rho \Rightarrow \begin{cases} \int \rho = \int \partial_i \partial_j E^{ij} = 0 \\ \int x^k \rho = \int x^k \partial_i \partial_j E^{ij} = - \int \partial_j E^{kj} = 0 \end{cases} ,$$

$\Rightarrow$  Global charge &  
dipole moment conservation



## Fractonic global field theories ?

- A signature for fractons: conservation of higher momentum.
- $\exists$  classification of the symmetries leading to it. [Gromov, PRX. '19]
- Simplest case:  $\phi(x) \rightarrow \phi(x) + \alpha + \beta_i x^i$ .

The Noether charges:

$$Q \equiv \int \rho, \quad Q^i \equiv \int x^i \rho, \quad \text{with } \rho = j_{(\alpha)}^0.$$

- Invariant theories: shift sym. with high derivative terms  $\rightarrow$  NG modes from space translation SSB excellent candidates!

Which leads us to...

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# What do we want to do ?

- SSB of spacetime symmetries  $\longrightarrow$  complicated  $\longrightarrow$  toy models.
- Concomitant breaking of translation and dilatation.
  - Condensed matter ( $\rightsquigarrow$  crystals).
  - Stat. physics (around critical points).
- Simplest case: homogeneous breaking  $\longrightarrow$  need internal  $U(1)$  SSB (+ permits a  $\mu$ ).
- Are there fractons somewhere ?

# The model

- In  $2 + 1$  d,  $\Phi$  a global- $U(1)$  complex scalar field &  $\Xi$  a real scalar field.

$$\mathcal{L} = \partial_t \Phi^* \partial_t \Phi + A \partial_i \Phi^* \partial_i \Phi + \frac{1}{2} \partial_t \Xi \partial_t \Xi - \frac{1}{2} \partial_i \Xi \partial_i \Xi - B \frac{(\partial_i \Phi^* \partial_i \Phi)^2}{\Xi^6} - H \Xi^6,$$

with  $A, B > 0$ , and  $H = \frac{A^2}{4B} \rightsquigarrow$  permits the energy minimisation.

- W/ natural units, notice that

$$[\Phi] = \frac{1}{2}, \quad [\Xi] = \frac{1}{2}, \quad [A] = [B] = [H] = 0.$$

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- Spatially-dependent stable background & homogeneous SSB  
 $\rightarrow$  2 kinds:

$$\left( \Phi_0 = \rho e^{ik_i x_i}, \Xi_0 = v \right) \text{ or } \left( \Phi_0 = b_i x_i, \Xi_0 = v \right).$$

- Approx. same qualitative results  $\rightarrow$  we discuss here

$$\Phi_0(t, x, y) = \rho e^{ikx},$$

$$\Xi_0(t, x, y) = v.$$

$$\mathcal{L} = \partial_t \Phi^* \partial_t \Phi + A \partial_i \Phi^* \partial_i \Phi + \frac{1}{2} \partial_t \Xi \partial_t \Xi - \frac{1}{2} \partial_i \Xi \partial_i \Xi - B \frac{(\partial_i \Phi^* \partial_i \Phi)^2}{\Xi^6} - H \Xi^6 ,$$

- $(\Phi_0 = \rho e^{ikx}, \Xi_0 = v)$  is a solution if  $\frac{k^2 \rho^2}{v^6} = \frac{A}{2B}$ .

$\rightsquigarrow$  2 flat directions  $\rightsquigarrow$  we expect 2 NG modes.

- Broken symmetries:  $P_1, R, D, S_R, S_I, Q$ .
- But:  $P_1 - kQ$  is an unbroken translation sym.  $\rightsquigarrow n_{\text{BS}} = 5$ .
- Considering the IHCs:

$$\begin{aligned} [P_1 - kQ, D] &\propto P_1, & [P_2, R] &\propto P_1, \\ [P_1 - kQ, S_R] &\propto S_I, & [P_1 - kQ, S_I] &\propto S_R. \end{aligned}$$

$\rightsquigarrow n_{\text{BS}}^{(\text{in})} = 2$ .

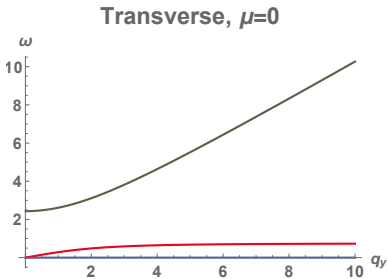
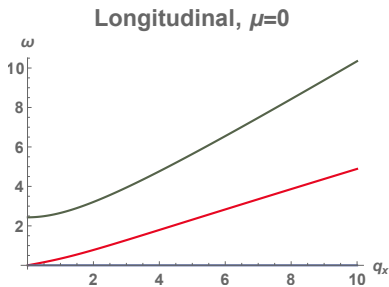


# Perturbative computation

- Small fluctuations around the background:

$$\Phi(t, x, y) = \rho e^{ikx} [1 + \sigma(t, x, y) + i\chi(t, x, y)] ,$$
$$\Xi(t, x, y) = v [1 + \tau(t, x, y)] .$$

- Abusive nomenclature:  $\chi$  the  $U(1)$ -mode,  $\sigma$  the shifton &  $\tau$  the dilaton. Consistent w/ Ward Ids. (cf. later).
- Numerics:



$$A = 0.125 , B = 0.25 , k = 1.5 , \rho = 1.07$$

# Ward-Takahashi Identities

- 3 Ward. Ids.  $\sim$  3 EOMs but when  $q \gg k$ , un-mixing  $\rightsquigarrow$  useful for modes identification.
- $U(1)$  Ward Id. at linear order:

$$\begin{aligned}\partial_t^2 \chi - 2A \partial_x^2 \chi &\simeq 0, \quad \text{with } k \rightarrow 0, \\ \Rightarrow \omega^2 &\simeq 2A q_x^2, \quad q \gg k.\end{aligned}$$

- Real shift Ward Id. at linear order:

$$\begin{aligned}\partial_t^2 \sigma &\simeq 0, \quad \text{with } k \rightarrow 0, \\ \Rightarrow \omega^2 &\simeq 0, \quad q \gg k.\end{aligned}$$

- Dilatation Ward Id. at linear order (based on improved  $T_{\mu\nu}$ ):

$$\begin{aligned}\partial_i \partial^i \tau - \partial_t^2 \tau &\simeq 0, \quad \text{with } k \rightarrow 0, \\ \Rightarrow \omega^2 &\simeq q_x^2 + q_y^2, \quad q \gg k.\end{aligned}$$

# Connection to fractons

- Quadratic action in the fluctuations:

$$\mathcal{L}_{\text{quad}} = \frac{v^2}{2} \partial_\mu \tau \partial^\mu \tau + \rho^2 (\partial_t \chi)^2 + \rho^2 (\partial_t \sigma)^2 - 2A\rho^2 [\partial_x \chi + k(\sigma - 3\tau)]^2.$$

- Emergent coordinate-dependent shift symmetries:

$$\delta\chi = \alpha(y) + \beta(x, y), \quad \delta\sigma = -\frac{1}{k} \partial_x \beta(x, y) + 3\delta + 3\gamma_i x^i, \quad \delta\tau(x, y) = \delta + \gamma_i x^i,$$

where  $\beta(x, y)$  &  $\alpha(y)$  are arbitrary functions.

- $\beta(x, y) \rightarrow$  immobile mode (fracton)  
 $\alpha(y) \rightarrow$  1-dimensional submode (lineon).
- At higher order, we lose this enhanced sym.  $\rightarrow$  fractonic modes need interactions to propagate.

$$\mathcal{L}_{\text{quad}} = \frac{v^2}{2} \partial_\mu \tau \partial^\mu \tau + \rho^2 (\partial_t \chi)^2 + \rho^2 (\partial_t \sigma)^2 - 2A\rho^2 [\partial_x \chi + k(\sigma - 3\tau)]^2 .$$

Toward the effective theory:

- Diagonalization of the mass terms:

$$\begin{pmatrix} v\tau \\ \sqrt{2}\rho\sigma \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \eta \\ \varphi \end{pmatrix}, \text{ with } \tan\theta = \frac{v}{3\sqrt{2}\rho}. \quad (*)$$

- $\eta$  is the massive mode with  $m_\eta^2 = 2Ak^2 \left(1 + 18\frac{\rho^2}{v^2}\right)$ .
- From (\*),  $v \gg \rho \Rightarrow \sigma \sim \eta$  and  $v \ll \rho \Rightarrow \tau \sim \eta$ .

# Connection to fractons

- By integrating out the massive mode  $\eta$  (low energy expansion):

$$\begin{aligned} \mathcal{L}_{\text{quad, eff}} = & \rho^2 (\partial_t \chi)^2 + \frac{1}{2} (\partial_t \varphi)^2 - \frac{1}{2} \sin^2 \theta (\partial_i \varphi)^2 + \frac{2\sqrt{A}\rho}{m_\eta} \sin \theta \cos \theta \partial_x \chi \partial_i^2 \varphi \\ & + \frac{2A\rho^2}{m_\eta^2} [(\partial_t \partial_x \chi)^2 - \cos^2 \theta (\partial_i \partial_x \chi)^2] + \frac{\sin^2 \theta \cos^2 \theta}{2m_\eta^2} (\partial_i^2 \varphi)^2 . \end{aligned}$$

- One of the symmetries:

$$\chi \rightarrow \chi + e + a_i x^i + c_{ij} x^i x^j + f(y) .$$

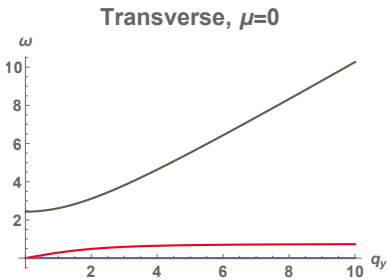
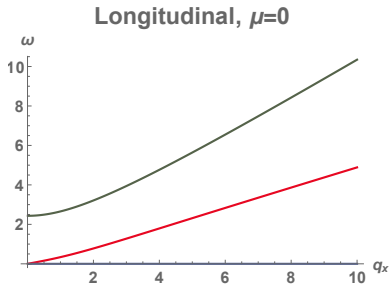
- Constant shift  $\rightarrow$  conserved charge.

Polynomial shift  $\rightarrow$  higher moment charge conservation.

$\rightsquigarrow$  signature of  $\chi$  as the immobile mode (at low energy).

$$\omega_\chi^2 \simeq 0 , \quad \omega_\varphi^2 \simeq \frac{v^2}{18\rho^2 + v^2} q_i^2 .$$

# Back to the numerics



$$A = 0.125, B = 0.250, k = 1.50, \rho = 1.07$$

# Complete dispersion relations

From the quadratic Lagrangian itself:

Compared to the W.Ids. and the EFT, we get

- $\omega_3$  at low energy.
- Subtlety at high momenta ( $q \gg m_\eta$ ):

$$\omega_2^2 \simeq \begin{cases} q_x^2 + q_y^2 & \text{if } (2A - 1)q_x^2 - q_y^2 > 0 \\ 2Aq_x^2 & \text{if } (2A - 1)q_x^2 - q_y^2 < 0 \end{cases},$$

and

$$\omega_3^2 \simeq \begin{cases} 2Aq_x^2 & \text{if } (2A - 1)q_x^2 - q_y^2 > 0 \\ q_x^2 + q_y^2 & \text{if } (2A - 1)q_x^2 - q_y^2 < 0 \end{cases}.$$

If  $A > \frac{1}{2}$ ,  $\omega_2$  and  $\omega_3$  swap their behaviour according to the direction in the  $(q_x, q_y)$  plane.

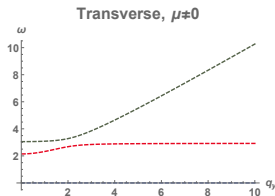
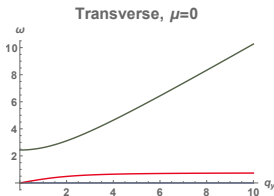
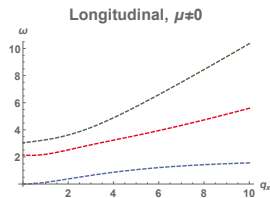
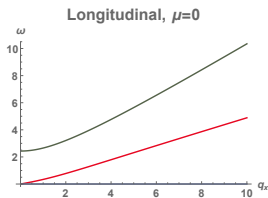
⇒ The **identification of the NG candidates** with the modes diagonalizing the kinetic matrix **depends on the norm and the direction of  $\vec{q}$** .

# A word about $\mu \neq 0$ case

- Small fluctuations around the background:

$$\Phi = \rho e^{i(\mu t + kx)} [1 + \sigma + i\chi] , \Xi = v [1 + \tau] .$$

- Numerics:  $A = 0.125$  ,  $B = 0.25$  ,  $k = 1.5$  ,  $\mu = 1$  ,  $\lambda = 0.5$



- Analytically:  $m_2^2 \sim \mu^{8/3} k^{-2/3}$  .



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## Fracton perspective:

- NG modes transforms “shift like”.
- SSB of translation requires higher derivatives.
- Associated EFTs are good candidates to display polynomial shift  $\rightsquigarrow$  fractonic signature!
- Indeed, from a **non-fractonic toy model, fractons emerged!**

## Goldstone perspective:

- Inverse Higgs Constrains have to be considered.
- Highly non-trivial modes identification.
- We have a gap scaling with  $\mu$  but “ugly”  $\neq$  internal case.

- Generalization of the toy model  
(has already been partially done in the paper).
- Exotic cases  $\rightsquigarrow$  holography.
- Phenomenological models similar to our toy model ?

Thank you!

# Switching on a chemical potential

## Why ?

- Conceptually interesting (comparison w/ internal sym.).
- To get closer to standard lab. conditions.

## How to handle a $\mu$ :

- At thermal equilibrium:

$$Z = \text{Tr} \left[ e^{-\beta(H - \mu Q)} \right] .$$

where  $\mu$  is the chem. pot. associated to  $U(1)$ .

- At  $T = 0$  ( $\beta \rightarrow +\infty$ ), the thermal vac. satisfies:

$$(H - \mu Q) |\mu\rangle = 0 .$$

- If  $U(1)$  SSB, then  $|\mu\rangle \sim e^{i\mu t}$  .
- $\Rightarrow$  in  $\mathcal{L}$ , term  $\sim \mu^2 \Rightarrow$  add. term to stabilise.

$$\mathcal{L} = \partial_t \Phi^* \partial_t \Phi + A \partial_i \Phi^* \partial_i \Phi + \frac{1}{2} \partial_t \Xi \partial_t \Xi - \frac{1}{2} \partial_i \Xi \partial_i \Xi - B \frac{(\partial_i \Phi^* \partial_i \Phi)^2}{\Xi^6} - H \Xi^6 - \lambda^2 (\Phi^* \Phi)^3 ,$$

- $(\Phi_0 = \rho e^{i(\mu t + kx)}, \Xi_0 = v)$  is a solution if  $\frac{k^2 \rho^2}{v^6} = \frac{A}{2B}$  &  $\mu^2 = 3\rho^4 \lambda^2$ .

$\rightsquigarrow$  1 flat direction  $\rightsquigarrow$  we expect 1 NG modes.

- Broken symmetries:  $P_0, P_1, R, D, Q$ .
- But:  $P_1 - kQ$  &  $P_0 - \mu Q$  are unbroken translation sym.  $\rightsquigarrow n_{\text{BS}} = 3$ .
- Considering the IHCs:

$$[P_1 - kQ, D] \propto P_1, \quad [P_2, R] \propto P_1,$$

$\rightsquigarrow n_{\text{BS}}^{(\text{in})} = 1$ .

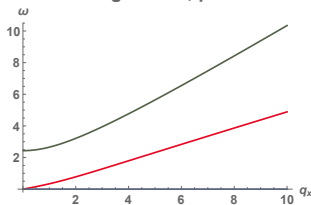
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- Small fluctuations around the background:

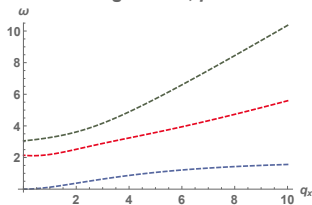
$$\Phi = \rho e^{i(\mu t + kx)} [1 + \sigma + i\chi] , \Xi = v [1 + \tau] .$$

- Numerics:  $A = 0.125$  ,  $B = 0.25$  ,  $k = 1.5$  ,  $\mu = 1$  ,  $\lambda = 0.5$

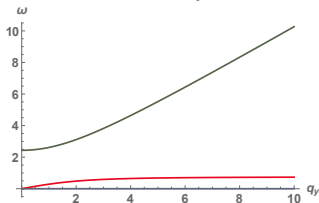
Longitudinal,  $\mu=0$



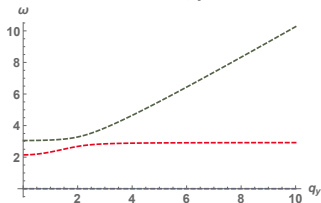
Longitudinal,  $\mu \neq 0$



Transverse,  $\mu=0$



Transverse,  $\mu \neq 0$



# Spectrum discussion

- Rem: analytics highly involved  $\rightarrow$  only parts of the spectrum.

## Some highlights:

- Enhanced shift sym. at quadratic order is lost.
- For  $\omega, k, \mu \ll q_x$ ,  $\omega_1^2 \approx 4\mu^2$ .
- For  $\mu \ll k$ ,  $m_3^2 = m_\eta^2$ ,  $m_2^2 \sim \mu^{8/3} k^{-2/3}$   
 $\rightsquigarrow$  “ugly” scale w/.  $\mu$ ,  $\neq$  internal case.
- Smooth limit  $\mu \rightarrow 0$ ,  $\lambda \rightarrow 0$  with  $\mu/\lambda = \text{cst.}$  reproduces  $\mu = 0$  case.  
 $\rightsquigarrow$  when  $\mu \ll k$  and/or  $\mu \ll q$ , the results are qualit. the same as  $\mu = 0$ .
- $\mu \gg k \gg q$  has been dealt w/.  $\rightsquigarrow$  confirm a hierarchy between  $m_3$  &  $m_2$ .